Recent developments at TERAPIX

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Outline

- TERAPIX and the CHTLS
- Access to software and support
- Development plan
- New developments for T04
- Ongoing software developments (for T05)



The team at IAP

- 4 astronomers:
 - Y.Mellier (PI/PM),
 E.Bertin (PS/PM),
 H.J.McCracken
 M.Schultheis
 (Besançon)
- 4 engineers:
 - M.Dantel-Fort,
 F.Magnard,
 C.Marmo,
 G.Semah
 - (Laurent Domisse left in April)
- 1 Ph.D student:
 - A.Baillard (EFIGI)





The tasks of TERAPIX

- Develop and distribute software tools required for the processing of MEGACAM and WIRCAM data
 - Pre-processing done at CFHT (Elixir)
 - TERAPIX software is developed in-house and can be used on various kinds of data
 - Released as Open Source to the community
- Produce and release calibrated, resampled, co-added images, weight maps and catalogs on a regular basis.
 - The 4th release is currently being processed
 - The achieved re-processing cycle time is about 12 months
 - Each release benefits from
 - Extended coverage
 - Complete re-calibration with increasing overlaps
 - Software upgrades and new features
- Provide support to members of the Canadian and French communities
 - Process P.I programs on request
- Manage data and hardware
 - Compute/storage farm of 26 bi-, quad-, and octo-procs for processing
 - 250 Gflops peak
 - Direct access to the data with 120TB of redundant storage
 - Cluster of 8 bi-pros available on request for data-intensive CFHTLS science





Access to software

- What is distributed:
 - source packages
 - Binary, static Linux RPMs (both x86 32 et 64 bits, single and multi-threaded)
 - PDF documentation
- Public unstable versions:
 - accessible through our SubVersioN repository
 - http://terapix.iap.fr/wsvn
 - "on demand" re-packaging of intermediary versions





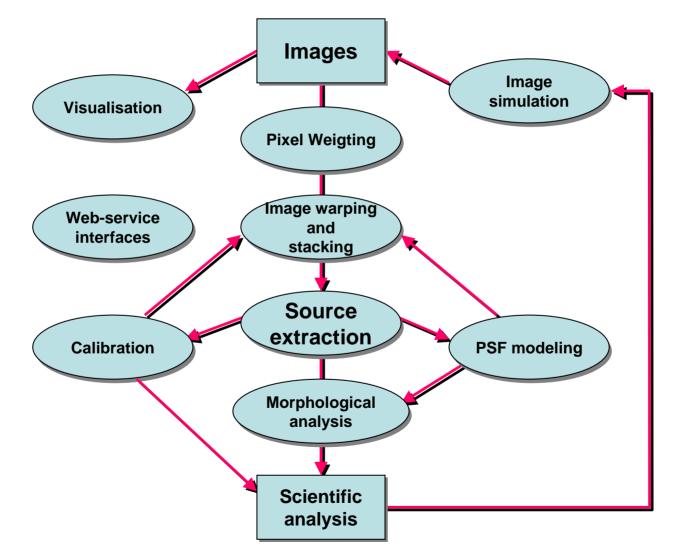
Getting help

- Discussion forums
 - http://terapix.iap.fr/forum
 - MEGACAM
 - CFHTLS
 - P.I.
 - WIRCAM
 - TERAPIX software
 - software for astronomy
 - hardware for astronomy
 - private forums
 - We can host your forum!

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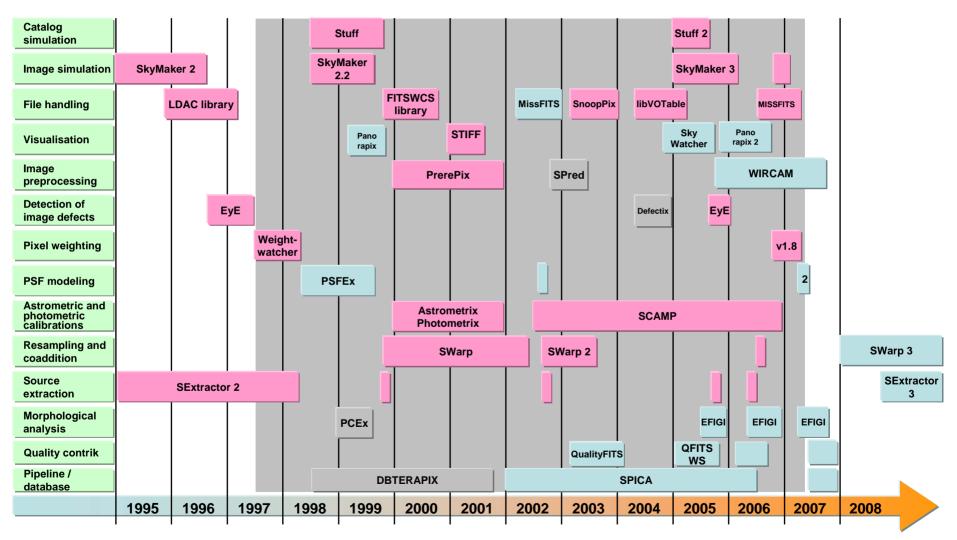


TERAPIX: An automated image analysis system





Development plan

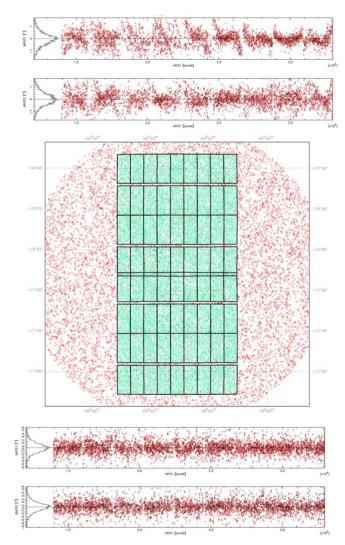


E.Bertin



New developments for T04

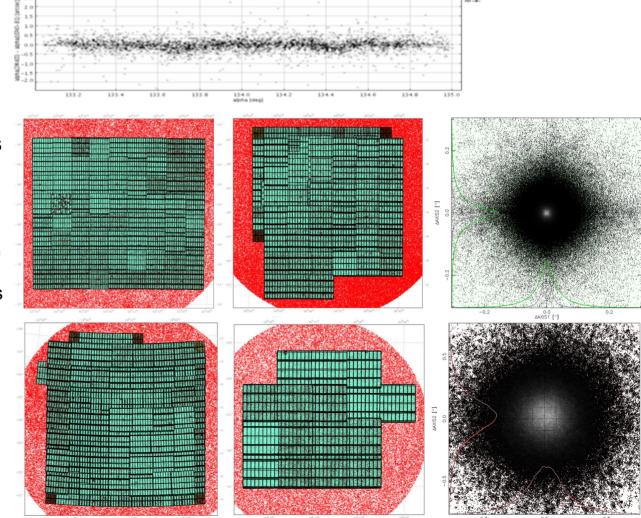
- SCAMP (astrometric and photometric calibration software): improve robustness of the existing approach
 - public release in 2006: testing by other users helped find bugs and improve algorithms
 - V1.2.1->1.3.8: recipes tuned to offer more robust behaviour in crowded fields and observing programs with poor dithering patterns.
 - WIRCAM processing (C.Marmo): very large number of exposures
 - used in various data challenges around the world
 - balancing between internal and external positional constraints
 - move to ICRS (at last)
 - Redistribute free parameters as the number of exposures increases





New developments for T04 (cont.)

- Astrometric reference changed from USNO-B1 to 2MASS
- Astrometric stats in T04 - Wide:
 - Up to 34.10⁶ sources used for solution and 48 instrumental "contexts"
 - Dispersion: (pairwise,3σ-clipped)
 - 27mas RMS internal for sources with S/N > 100
 - 0.23" RMS with respect to 2MASS
 - Systematics with respect to 2MASS ≤0.13"



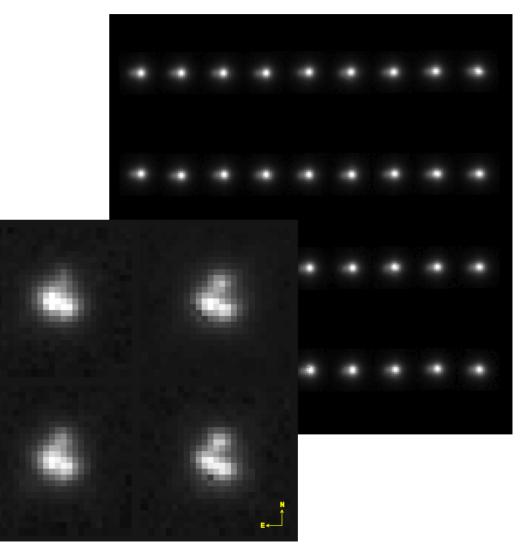
CFHTLS users' meeting 05/2007

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Ongoing software developments for T05

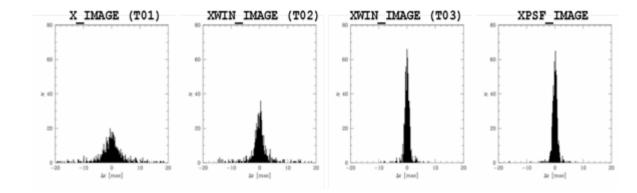
- Automated quality control
 - New PSFEx version
 - Will be released later this year
 - New metrics to track down bad PSFs (defocused, trailed, multimodal)
 - Works with undersampled data





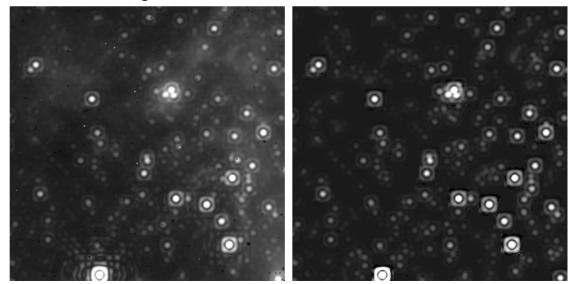
Ongoing software developments for T05 (cont.)

- PSF-fitting photometry
 - Mostly improve photometry
 - Improvements and testing contributed by Ph. Delorme (Grenoble)
 - Deblending issue:
 Extragalactic
 science vs galactic
 science



original

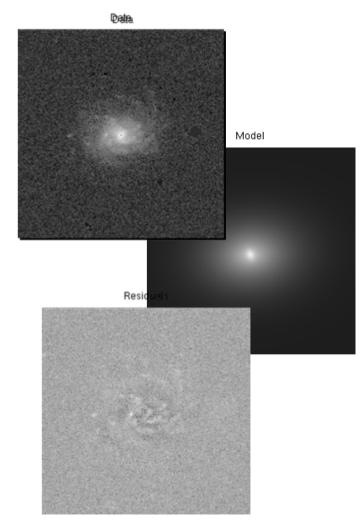
reconstructed





Ongoing software developments for T05 (cont.)

- Galaxy profile-fitting
 - PSF model available
 - varies smoothly over Deep and Wide stacks
 - Development done in the framework of the EFIGI project
 - Analytical profiles
 - Single Sersic
 - Sersic + exponential (12 free parameters)
 - About 10 (faint) galaxies per second per 2GHz core
 - Suitability for lensing studies needs to be assessed
 - Use for star-galaxy separation





Ongoing software developments for T05 (cont.)

- QualityFITS pipeline (automated pixel weighting / image diagnostic) moved to CFH
- New pipeline
 - Improved job balancing and data handling
 - Allow images from various imagers to be processed transparently
 - "shopping cart" model with full control on what is processed and what is not
 - Rely heavily on XML outputs from pipeline modules and the associated XSLT filters for visual checks and tracking of software configurations
 - More consistent use of flag maps



by user bertin from kiravix.lap.fr in /diak2/seasp/ceowded

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