The CFHT Legacy Survey

Operations and Services at CFHT

Jean-Charles Cuillandre - Canada-France-Hawaii Telescope

CFHTLS National Meeting - France - IAP/Terapix - Feb. 04, 2005









The MegaPrime Project

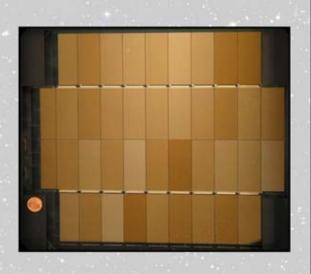
The components

- **Variable Variable V**
- √ Wide-field Corrector (HIA / SAGEM-REOSC) 4 lenses
- √ Image stabilizing unit (Obs. de Paris) Tip-tilt correction up to 5Hz
- √ Focus-guiding unit (HIA / CFHT) 2 guiders & autofocus
- √ MegaCam (CEA) 40 CCDs mosaic

The specifications

- √ 1 full square degree field of view
- √ Properly samples 0.7 arcsec. seeing
- √ Sensitive from the u to z band (Sloan filters)
- √ Image quality: 0.1 arcsec from center to edge
- √ Low overheads (detector limited)
- **♦** Operated within CFHT's New Observing Process (QSO + Elixir + DADS)







MegaPrime & MegaCam first 2 years

- **√** Complex instrument
- √ High usage: 55% of the telescope time
- √ Sensitive operational load: requires preventive maintenance
- √ Still some on-going development & tuning (image quality & autofocus)
- √ Despite difficulties, MegaPrime has had only a few nights of down time

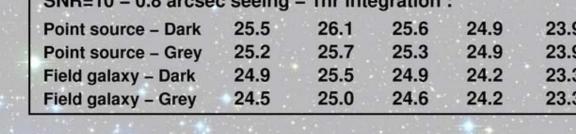
Time range	QSO run(s)	Event	Effect on data
02 Dec. 04 - present	04BQ05 and on	L3 lens flipped upside down	Spectacular IQ improvement
01 Nov 04 Nov. 04	04BQ04	L3 lens flipped upside down	Spectacular IQ improvement
08 Sep 22 Oct. 04	04BQ02-Q03	Defective video board	MAXLINB=32K right amp CCD27->35
05 Sep 07 Sep. 04	04BQ02	Defective video board	Half mosaic not operational (18-35)
14 Aug 24 Aug. 04	04BQ01	CCDs 00,11 defective	Only A amplifier (left) off for 07 and 11
06 Aug 14 Aug. 04	04BQ01	CCDs 00,07,09,10,11 defective	Only A amplifier (left) off for 00 and 11
16 Aug 17 Aug. 04	04BQ01	Ice on window dusk to 1 AM	Large blob obstructing light
06 Aug 07 Aug. 04	04BQ01	Ice on window all night	Large blob obstructing light
06 Aug 10 Aug. 04	04BQ01	Vacuum gauge glow	Large structures in all bands
07 Jul. 04 - present	04AQ07 and on	L1 cell O-ring removal	No obvious IQ improvement
25 Feb. 04 - present	04AQ01 and on	Light baffle installed	Cut stray lights, lower sky background
13 Jan 30 Jan. 03	03BQ07	CCD03 non operational	2.7% off the field of view at the edge
16 Dec 31 Dec. 03	03BQ06	CCD03 non operational	2.7% off the field of view at the edge
28 Nov 03 Dec. 03	03BQ05	CCD03 non operational	2.7% off the field of view at the edge
27 Oct 01 Nov. 03	03BQ04	CCD03 non operational	2.7% off the field of view at the edge
22 Oct 22 Oct. 03	03BQ04	Ice on window 3 AM to dawn	Large blob obstructing light
18 Sep 04 Oct. 03	03BQ03	CCD03 non operational	2.7% off the field of view at the edge
18 Jul. 03 - present	03AQ06-present	WFC original tilt & centering with 5.5mm spacer	Back to best optical setup
16 Jun 17 Jul. 03	03AQ05	WFC 0.02 deg. tilt to west side with 5.5mm spacer	Not optimal optical setup yet

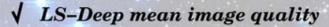


MegaPrime+ MegaCam Performance

- 350 million pixels (~18000x18000) 700 Mb per image
- Geometry: 0.96 x 0.94 sq. degree
- Sampling: 0.187 arcsec/pixel
- Improved image quality: only 2 corner CCDs with poor IQ
- Excellent immunity to scattered light
- Photometric performance:

Filter	u*	g'	r'	ľ	z'
SNR=10 - 0.8 arcse	c seeing	– 1hr inte	egration		
Point source - Dark	25.5	26.1	25.6	24.9	23.9
Point source - Grey	25.2	25.7	25.3	24.9	23.9
Field galaxy - Dark	24.9	25.5	24.9	24.2	23.3
Field galaxy - Grey	24.5	25.0	24.6	24.2	23.3





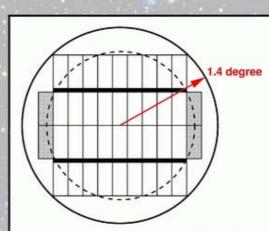
u* = 0.89 arcsec.

g' = 0.89 arcsec.

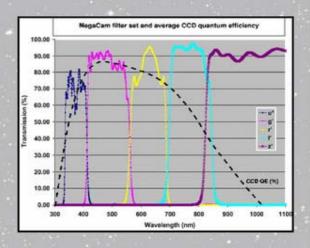
r' = 0.84 arcsec.

i' = 0.80 arcsec.

z' = 0.76 arcsec.







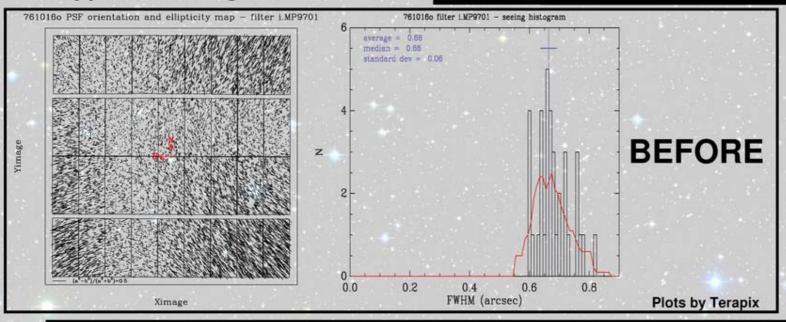


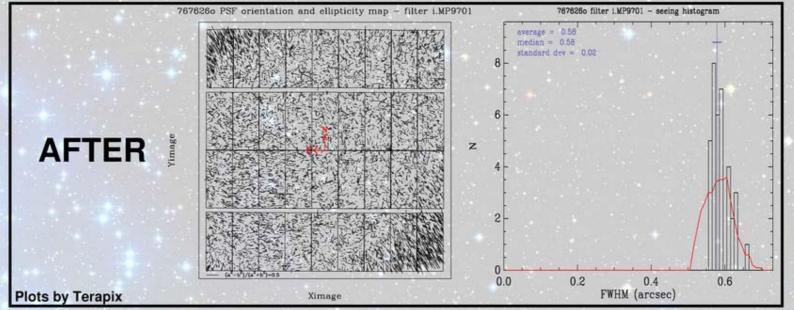
Dramatic IQ improvement

The flipped L3 configuration

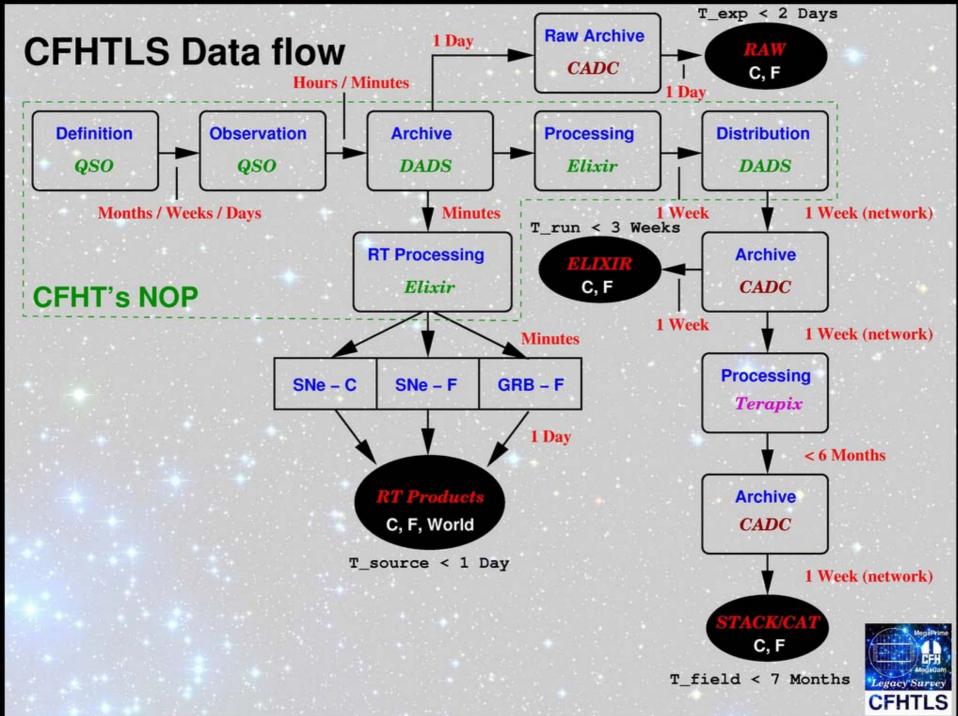


10.5410.4910.4810.4910.4710.4610.4610.4910.541 10.4910.5010.4910.4710.4610.4610.4510.4410.451 10.5110.5310.5410.5010.4710.4710.4610.4610.491 10.5310.5410.5410.5310.5010.4810.4810.5410.651









Services: entities serving the CFHT Legacy Survey community

CFHT

Gather, Detrend & Calibrate

- √ Get the raw data (QSO)
- Detrend the data (Elixir)
- √ Calibrate the data (Elixir)
- √ Collect the Meta–Data (DADS)
- Ship data to CADC (DADS)
- **√** Quality control
- √ Progress report

Terapix

Calibrate, Stack & Extract

√ Science images:

Fully calibrated

Stacked

- Source catalogs
- √ Quality control
- √ PI service

CADC

Archive & Distribute

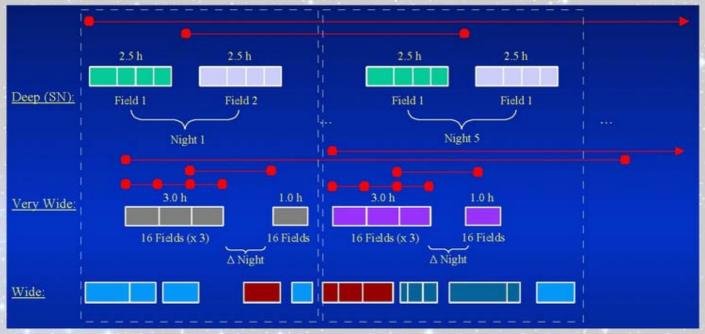
- √ Raw MegaCam FITS data
- √ CFHT data products
- √ Terapix data products

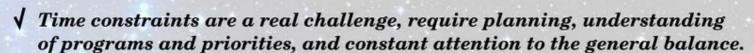


The CFHTLS seen by the Queued Servive Observing (QSO) A complex endeavor:

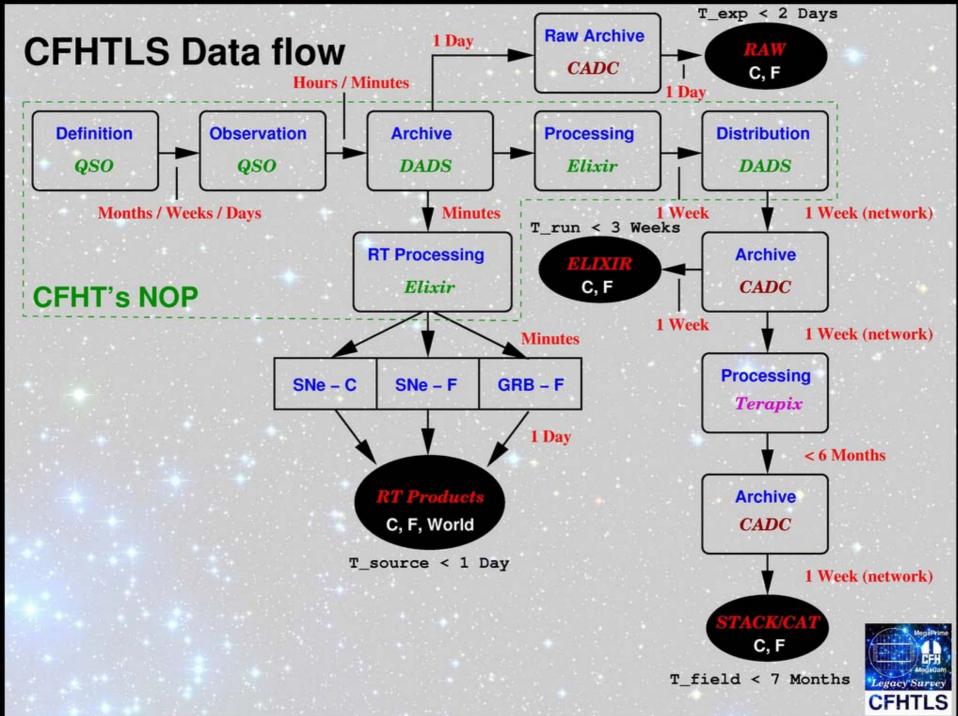
- √ Time critical observations on a large-scale:
 - SNe fields every 4 nights, 2 fields per run Each Very Wide patch has 5 time constraints
- **V** Balancing Agency Time (acces time not equal to guaranteed time!)
 - PI programs (50%) versus CFHTLS (50% or "49" nights at best per semester)
- √ Balancing time share between CFHTLS programs

 Deep (44%), Wide (34%), Very Wide (22%)
- **√** Example of conflicts just with the CFHTLS:



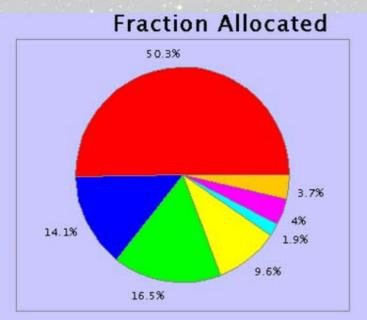


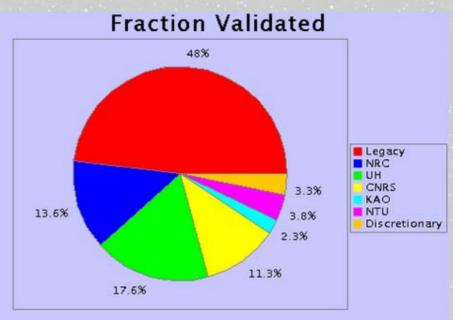




Balancing the agencies and the CFHTLS components

√ Agencies balancing (TOP priority): (for semester 04B)





- √ CFHTLS balancing: (as of Feb. 2005)
 - Requested: Deep/SNe = 44%, Wide = 34%, Very Wide = 22%
 - Validated: Deep/SNe = 58%, Wide = 25%, $Very\ Wide = 17\%$
- **▼** Balancing CFHTLS is difficult due to constraints, and is VERY weather dependant
- **V** Effective work by the Steering Group coordinators to establish priorities per run

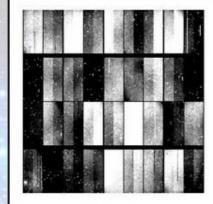


Elixir detrending & calibration

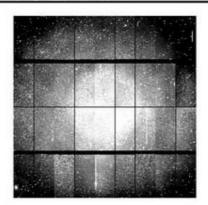
- **√** Multi-Extension FITS format exclusively (MEF)
- √ Data splicing: go from two readouts to a single detector configuration
- √ Pixels masking: less than 0.2% of the mosaic
- √ Overscan & bias: gradient in overscan and pixel ringing
- √ No dark current, hence no correction
- √ Flat-fielding: twilight flat-fields (has scattered light)
- √ Photometric superflat: 0.7% photometric flatness across the image
- √ Fringes: 6% in i', 15% in z' Residual < 1% To be refined!
 </p>
- √ Sky background modes Dismissed!
- √ Astrometric & photometric calibration To be refined!

Photometric Flat

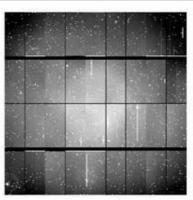
The detrending steps of Elixir (3 mn total):



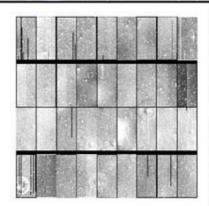
RAW



FLATTEN



DEFRINGE



DEMODES



Data & Computing facilities

√ Data so far (Feb. 2005): 42,000 frames archived = 30 Terabytes

√ Hardware:

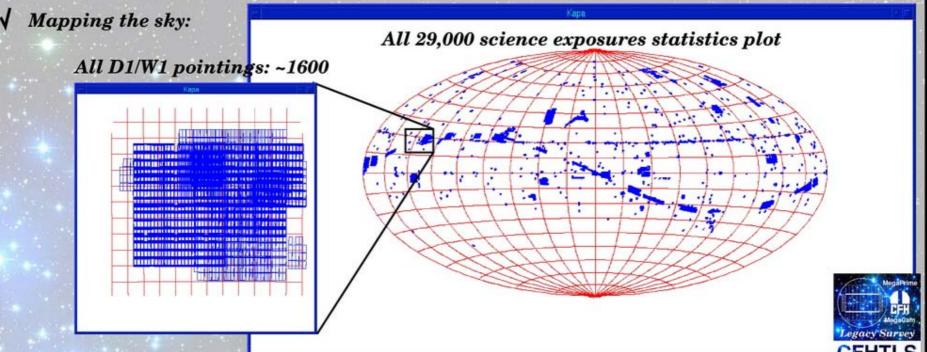
• Archive (DADS): 12 nodes / 15 Tbytes storage / 20 CPU = 31 GHz

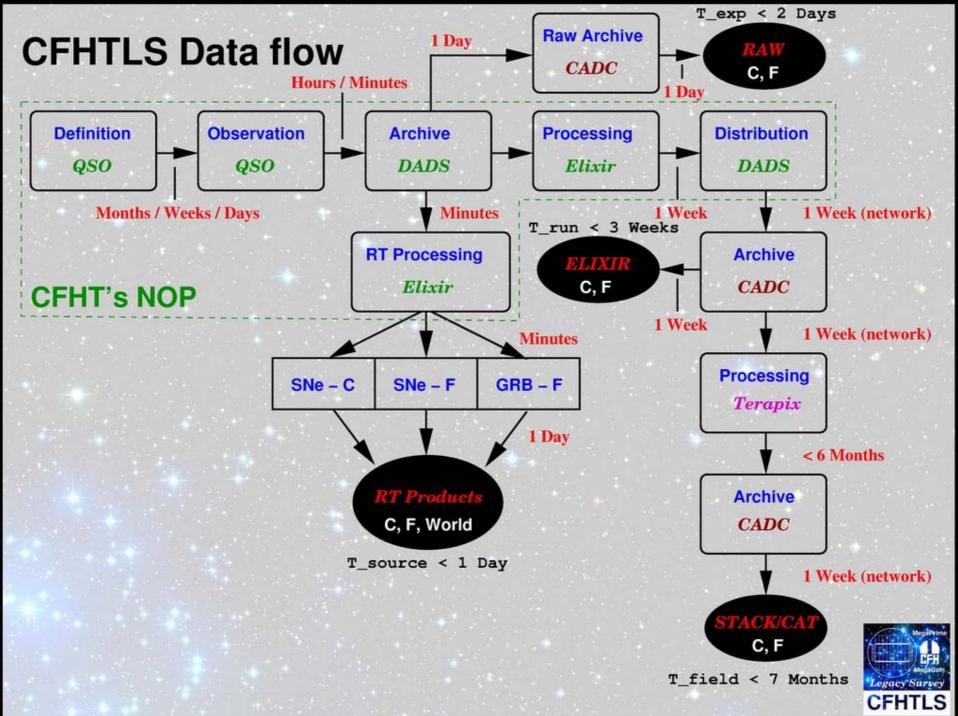
• Processing (Elixir): 13 nodes / 3 Tbytes storage / 20 CPU = 48 GHz

• Analysis (Real Time Systems): 9 nodes / 12 Tbytes storage / 18 CPU = 42 GHz

Total: 33 nodes / 30 Tbytes storage / 58 CPU = 1 THz

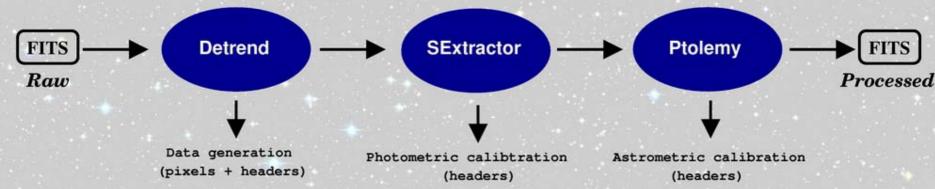
V People: DADS: 1.5 Elixir: 1



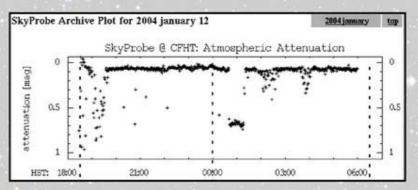


Elixir calibration & data products

√ Elixir pipeline:



- √ Astrometry: on a per CCD basis (0.2 arcsec. accuracy)
- √ Photometry: one zero point per filter per run based on numerous standard fields observations
- **√** Master detrend data (Mask, Bias, Flat-fields, Fringe, Modes)
- **√** Meta-Data tables:
 - Processed images information
 - Master detrend information
 - Standard stars images zero points
 - Standard star transparency
 - SkyProbe transparency
 - QSO comments data
 - Weather data
 - Supporting graphics images





CFHTLS: all Elixir data currently available from June 2003 to today

22 QSO Observing Runs = 6624 validated exposures for 472 hours of light integration

✓ Deep Survey: 58% (vs target 44%)

3044 exposures / 271 hours of light integration

D1: 946 exposures / 94 hours integration

D2: 408 exposures / 36 hours integration

D3: 720 exposures / 51 hours integration

D4: 901 exposures / 90 hours integration

Wide Survey: 25% (vs target 34%)

1036 exposures / 122 hours of light integration

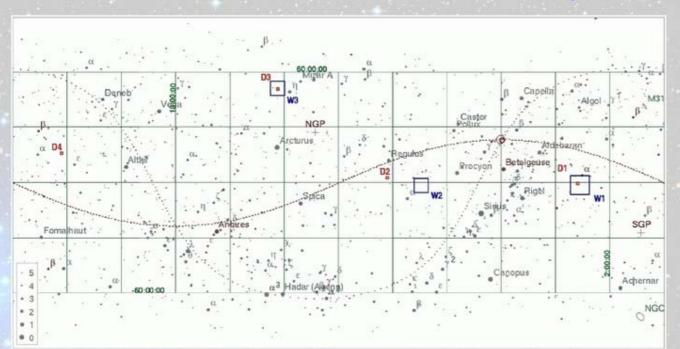
W1: 499 exposures / 75 hours integration

W2: 238 exposures / 23 hours integration

W3: 133 exposures / 17 hours integration

√ Very Wide Survey: 17% (vs target 22%)

2544 exposures / 79 hours of light integration



Mean image quality all Deep exposures

u* = 0.89 arcsec.

g' = 0.89 arcsec.

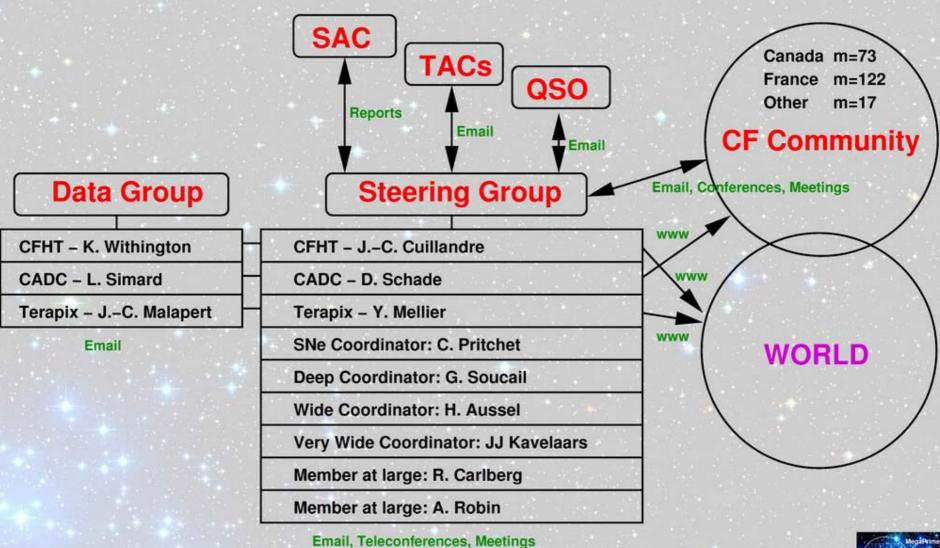
r' = 0.84 arcsec.

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The People & Communication channels





Realities of observing:

- **V** Bad weather account for *at least* 20% of time lost
- √ Technical problems accounts range from 20 (04A) to 6 nights (04B) a semester (10% avg)
- √ Validation process: typically 85%, 03B was only 80% due to poor seeing

Remaining means to increase the open shutter time per night

√ Overheads per night = 3 hours (typical length of a night: 9.5 hours)

	Expo	sure	over	heads	
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$$70 \times 50 \text{ sec} = 58 \text{ mn}$$
 Detector limited

$$30 \times 0.4 mn = 12 mn$$
 Optimized

$$20 \times 1.5 = 30 \, mn$$
 Difficult to reduce

$$6 x 7 = 42 mn$$
 Soon to be 0 mn (!?)

$$12 \times 2 = 24 \text{ mn}$$
 Could be 0 with tertiary stds.

... about 1 hour could still be gained.

All this contributes for the initial CFHTLS goals not to be achieved ... but things keep getting better each new semester!

 $5 \times 2 \, mn = 10 \, mn$



Advertising the CFHTLS in 2004 (mostly by SG members)

The CFHTLS Operation

CFHT: Services provided for the CFHTLS	
LSST Symposium	- Sep. 2004
SFZA French Astronomy national meeting	- Jun. 2004
CFHT Users' Meeting	- May. 2004
CFHTLS French Mational users meeting	- Jan. 2004
CADC: CFHTLS & Archiving	
Seminar in Toulouse	- Nov. 2004
LSST Symposium	- Sep. 2004
Cowichan Valley Star Finders	- Jul. 2004
HIA Science Coffee	- Jul. 2004
Astrometry in the Age of Large Telescope	- Jun. 2004
Canadian Astronomical Society	- Jun. 2004
BC Teachers Association	- Feb. 2004
CFHT Users' Meeting	- May. 2004
Terapix: CFHTLS & Data Processing	
The Annual Astrowise Meeting	- Dec. 2004
WIRCam workshop on Large Programs	- Nov. 2004
The AVO seminar in Meudon	- Nov. 2004
CFHT 25th Anniversary Press Release	- Sep. 2004
THE ADASS Meeting	- Sep. 2004
Seminars in Marseille, Besancon, Lyon	- Fall 2004
SFZA French Astronomy national meeting	- Jun. 2004
CFHT Users' Meeting	- May. 2004
The AVO meeting in Meudon	- Mar. 2004
General Public conference at Uranoscope	- Feb. 2004
Public Lecture at the College de France	- Feb. 2004
CFHTLS French National users meeting	- Jan. 2004

The CFHTLS Science

Very Wide: CFHTLS & Solar System	
Division of Planetary Science Meeting CFHT Users' Meeting skiLS Meeting Canada France Ecliptic Plane Survey	 Nov. 2004 May. 2004 Feb. 2004 Jan. 2004
Wide: CFHTLS & Lensing	
Shanghai Summer School in cosmology Cosmos04 Meeting in Toronto General Public conference at Bellem IAU Symp. Impact of lensing on cosmology Colloquium in Heidelberg	- Oct. 2004 - Sep. 2004 - Aug. 2004 - Jul. 2004 - Jun. 2004
SNLS: Supernovae and Dark Energy with th	e CFHTLS
SNLS 1st year results — Texas Meeting ESO seminar Canadian Undergraduate Physics Conferenc	- Dec. 2004 - Nov. 2004
Caltech seminar	- Oct. 2004
Cosmos04 Meeting in Toronto SF2A French Astronomy national meeting AAS Meeting Denver	- Sep. 2004 - Jun. 2004 - Jun. 2004
Padova Conference CFHT Users' Meeting	- Jun. 2004 - May. 2004
International Gemini Science Meeting Observing Dark Energy (NOAO Workshop)	
CIAR Cosmology and Gravitation Meeting AAS Meeting Atlanta	- Mar. 2004 - Jan. 2004
Rencontres de Moriond	- Jan. 2004



Facts, conclusions and perspectives

- MegaPrime / MegaCam is more and more reliable
- ✓ Image quality has (luckily) improved dramatically
- ✓ Instrument operation overheads too high, need to be further reduced
- ✓ Increased support from the coordinators to the QSO team effort
- **№** V Time constraints affect the overall survey image quality
- ② ✓ Delay in the first Terapix release due to recipes complexity (& Elixir first release in Jan.04)
- ✓ Elixir data available to the community at CADC within 2 to 3 weeks after a run
- Solution I Benefit to Terapix for regular steady releases to the LS community
- ✓ CFHTLS Registered Users download massive amounts of LS data from CADC
- Science cruising on the real-time components in well organized collaborations
- Steering Group has made a large effort on communications with the community
- \checkmark More interactions (reports) between the SAC, TACs and the SG would benefit the survey
 - **▼** CFHTLS Workshop to be held in Paris in May 2005 (both C&F communities)
 - **Vork already starting for the preparation of the CFHTLS review in spring 2005**

Legacy Survey
CFHTLS

www.cfht.hawaii.edu/LSW05/