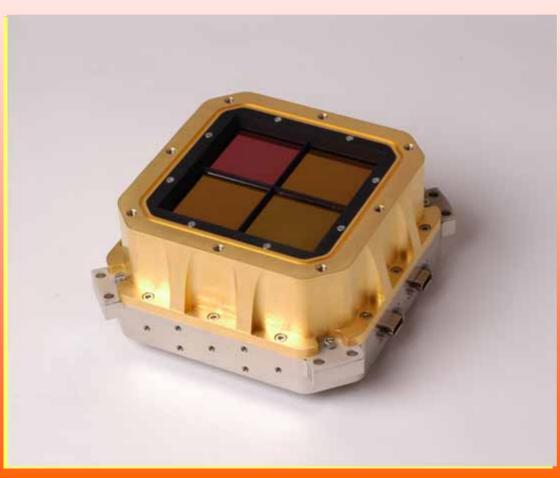


WIRCam data reduction @ TERAPIX



Chiara Marmo IAP 6-3-2006



Technical informations



Number of detectors	4 = 2 x 2
Detector size (pixel)	2040 x 2040 active pixels (2048 x 2048 with reference pixels)
Camera field of view	21.5 arcminute
Field distortion	<0.8% in the corners
Gaps between CCDs (arcsec)	45
Magnitude system	AB
Available filters	0



Processing steps

- image cubes splitting, if necessary;
- quality assessment and weight maps production;
- precise astrometric and photometric calibrations;
- □ stack generation;
- □ catalogs and final quality assessment delivery.



Quality assessment

Using QualityFITS:

bad pixel identification is made using an ad-hoc sextractor filter (builded using EyE), cosmic_Wircam.ret;

a .reg file is also introduced in building weight maps, masking edge zones strongly affected by bad pixels;

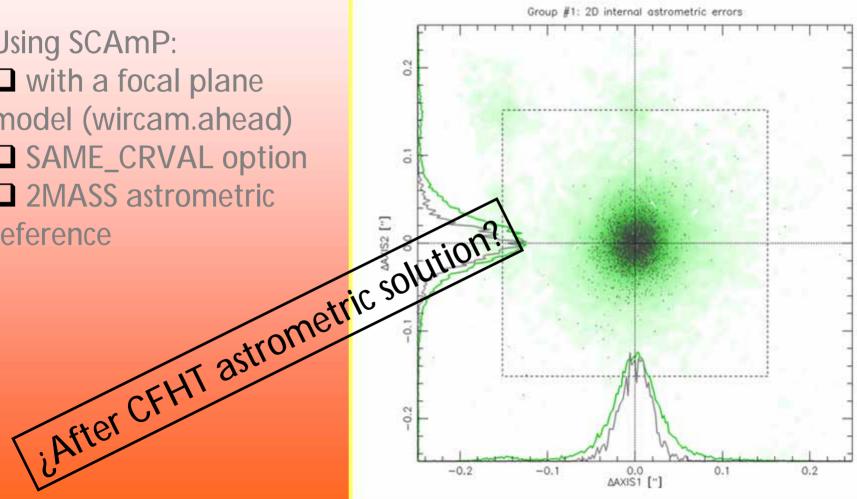
Sextractor parameters are optimised to eliminate spurious detections:

- ✓ DETECT_THRESH 4
- ✓ INTERP_TYPE NONE



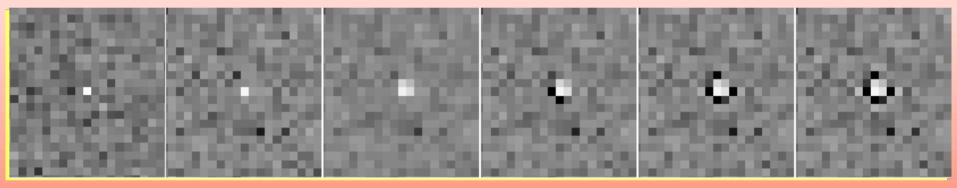
Astrometry

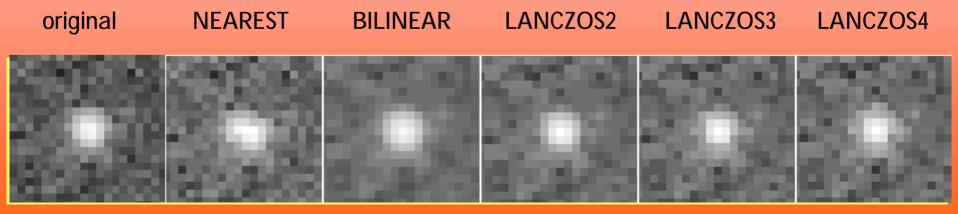
Using SCAmP: with a focal plane model (wircam.ahead) □ SAME_CRVAL option □ 2MASS astrometric reference





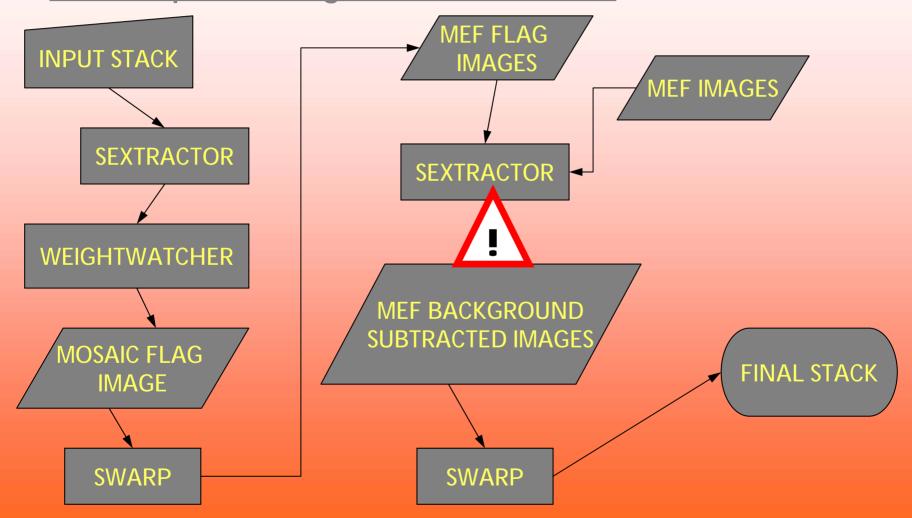
Stacking Using SWarp: RESAMPLING_TYPE LANCZOS2, in order to avoid under-sampling effects.





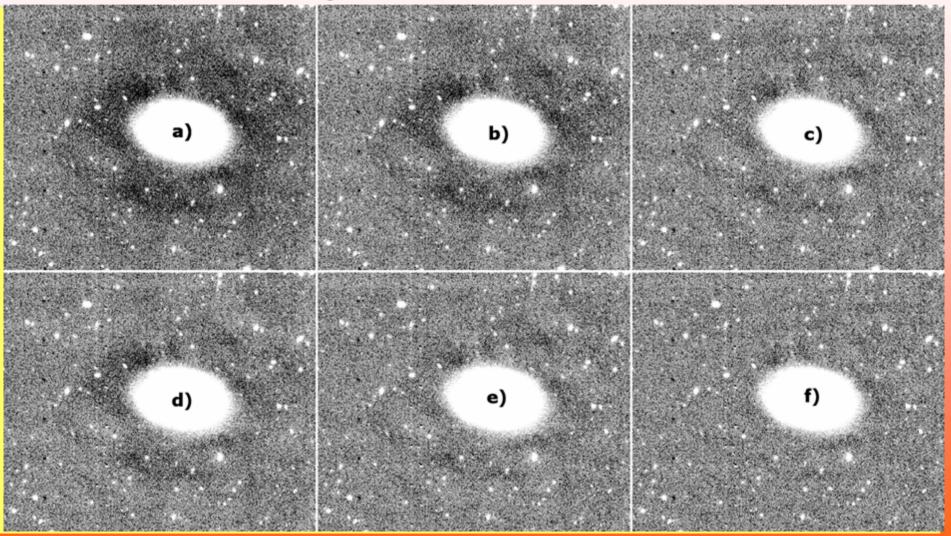


Double pass background subtraction





Double pass background subtraction





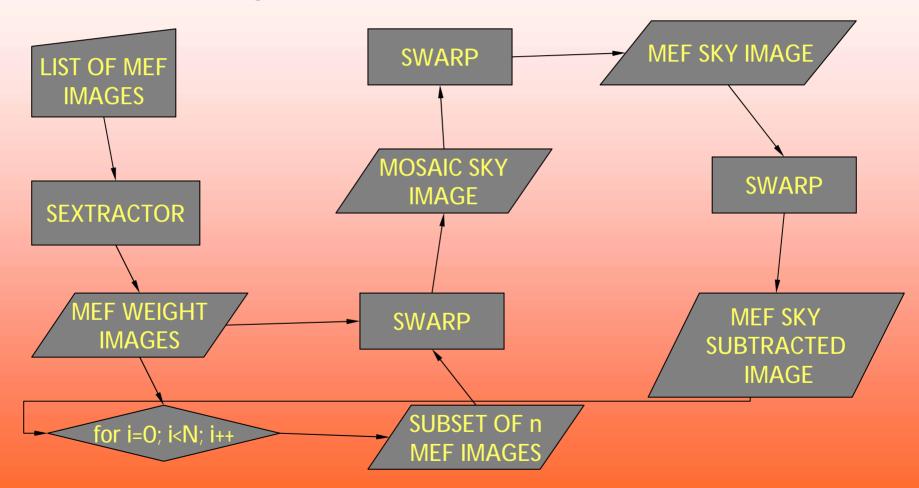
Final products

Using QualityFITS:

- seeing and background estimation
- preliminar star galaxy separation
- catalogue containing basic object parameters
- flag image used to flag objects in the catalogue
- \Box χ^2 image and merged catalogue
- •

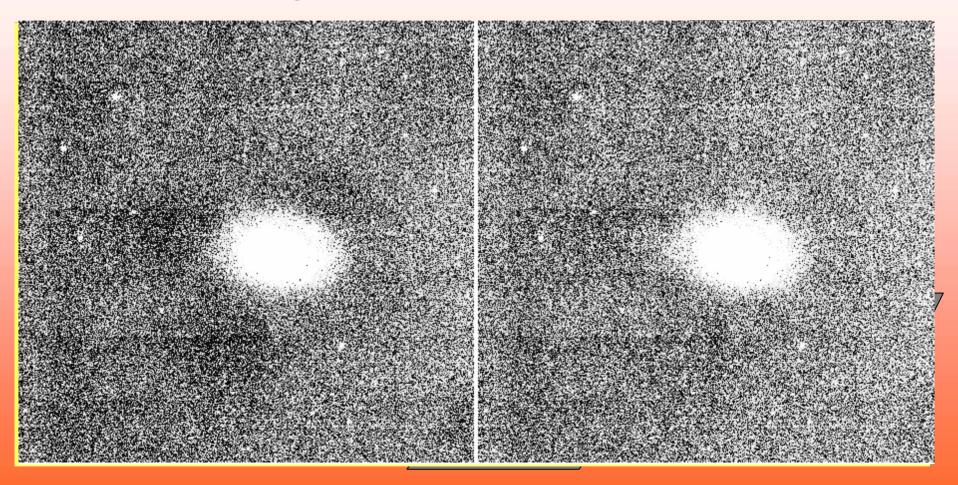


Extras: initial sky subtraction





Extras: initial sky subtraction

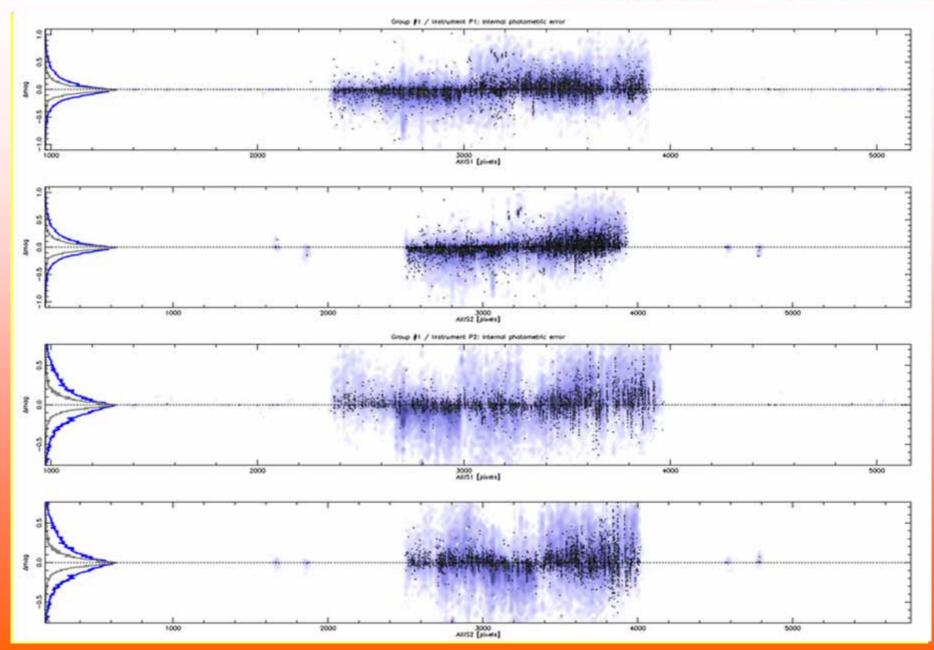




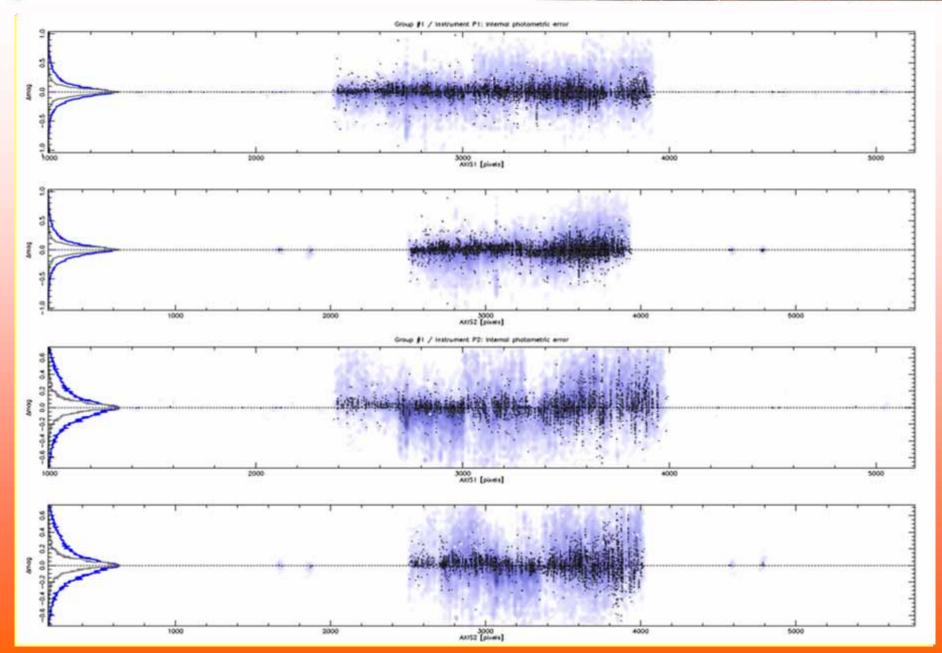
Extras: detector zero-point determination

- SCAmP provides the same rigid shift of photometric zero-point for all extensions in a MEF file.
- How to calibrate each detector separately?
- compute the astrometry as usual
- □ split the MEF files (and their headers)
- □ find one ore more photometric reference image or (2MASS?) and label it as photometric
- □ SOLVE_PHOTOM with SCAmP fixing the astrometry
- join the headers





N TERAPIX





Summary

What does WIRCam data reduction need from SPICA?

- □ To verify NAXIS keywords and split cubes, if necessary.
- □ To implement new qualityFITS functions.
- To introduce WIRCam configuration files or specific options for qualityFITS, Sextractor, Scamp, SWarp.
- To introduce an option for double pass background subtraction.

Open questions

- Database structure.
- $\hfill \label{eq:general}$ Final χ^2 images and flagged catalogs.